

The ANDES laboratory: Contributions from Seismology and Geophysics

Andreas Rietbrock (University of Liverpool, KIT)

&

Thomas Forbriger (Black Forest Observatory, KIT)

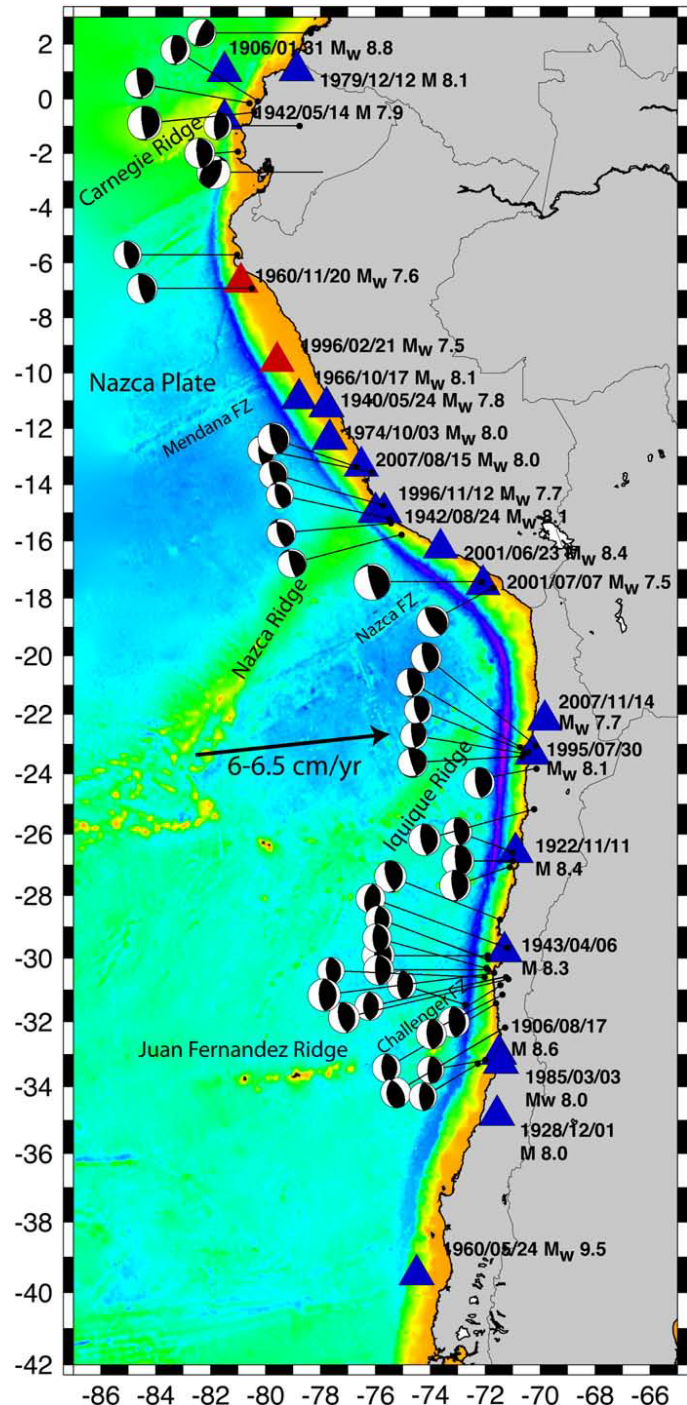
Outline

- Seismicity and modes of displacement in a subduction zone environments
- Underground observatory:
 Black Forest Observatory (BFO)
- Opportunities (Wish list...) & suggestions

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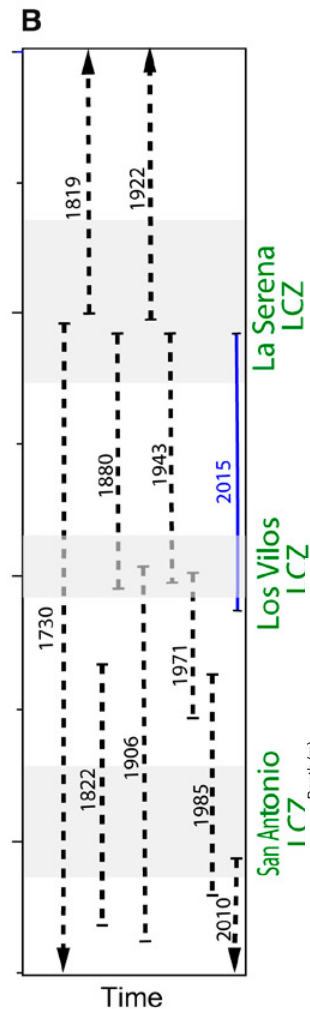
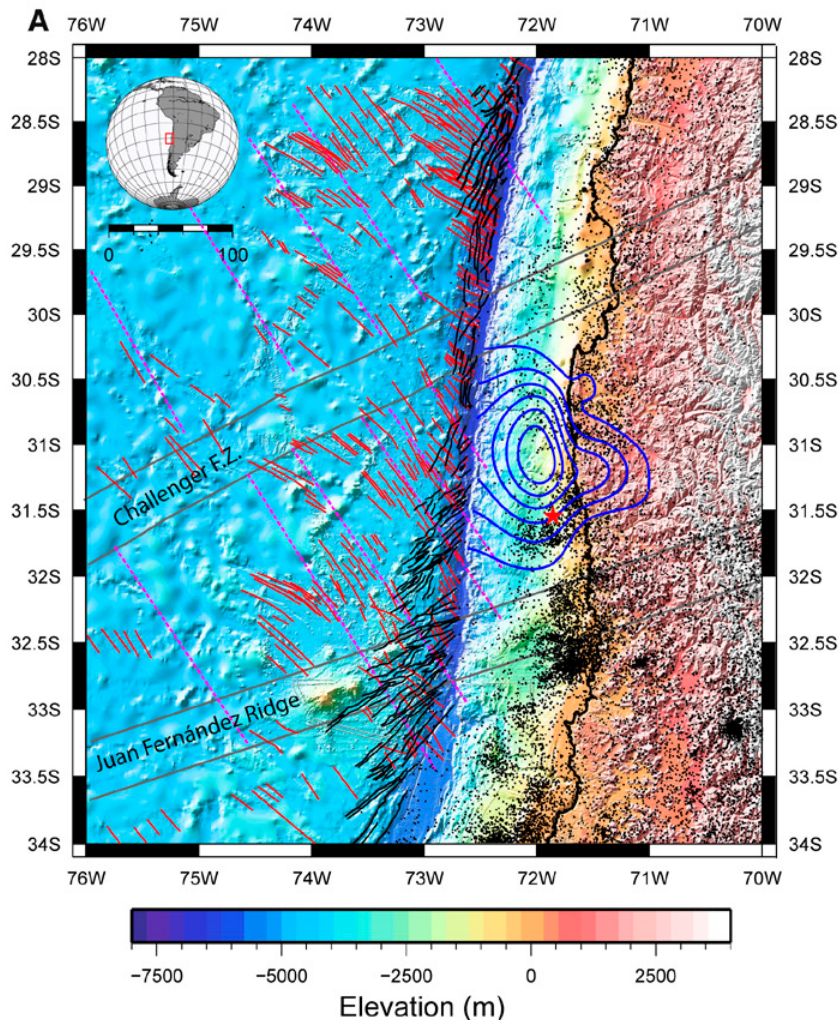
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South American Subduction Zone

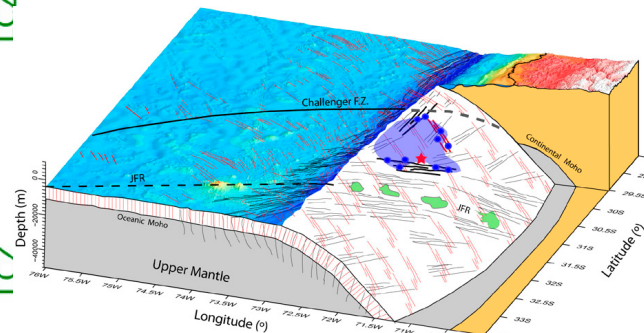


- Ocean-Continent collision
- Nazca plates subducts beneath the South American continent
- 2nd highest mountain range in the world
- Largest ever instrumentally recorded earthquake:
1960 Valdivia earthquake M_w 9.5

The 2015 Mw 8.3 Illapel earthquake

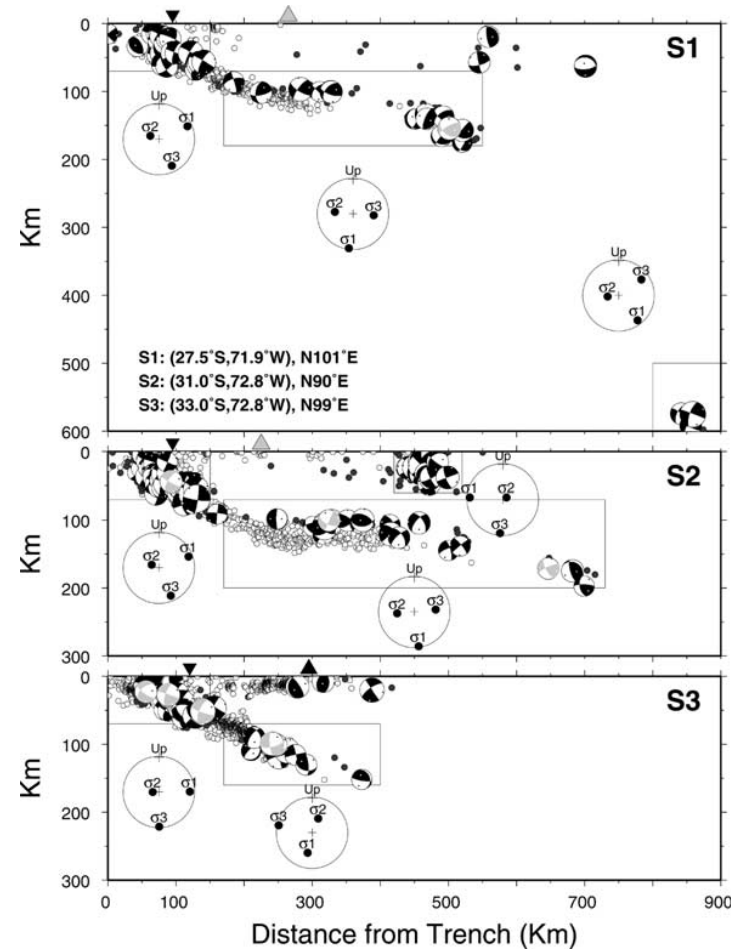
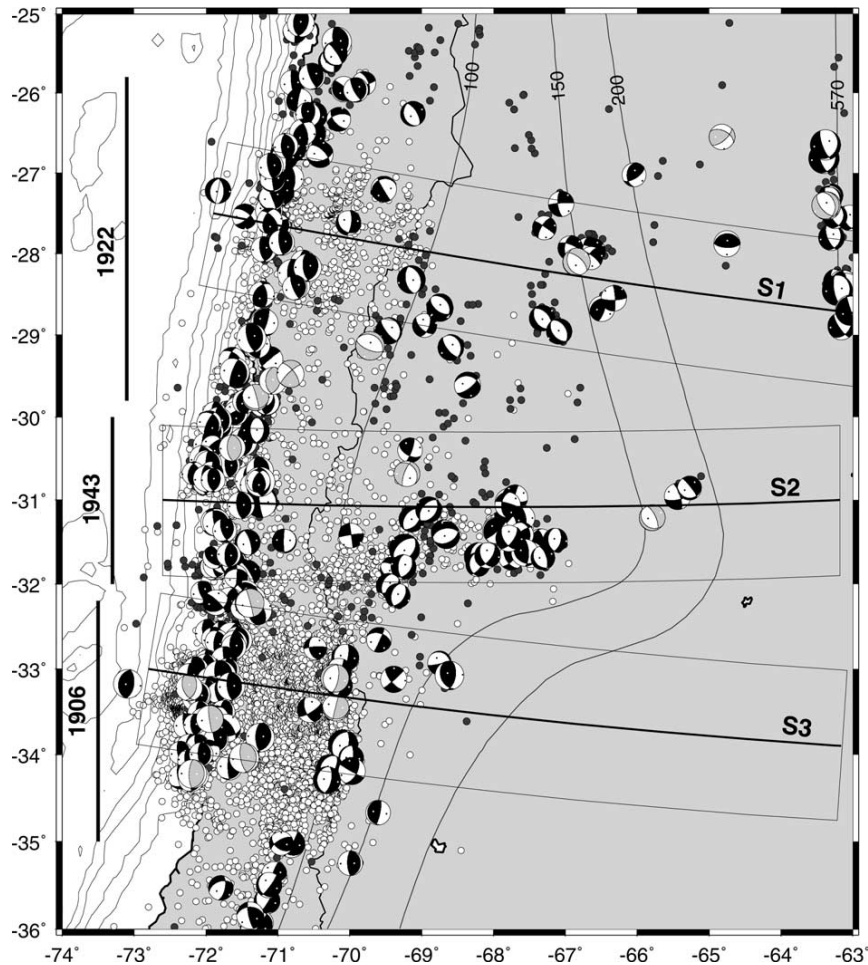


- Good historical record
- Large earthquake occur at “regular” intervals
- Fracture zones seem to limit the rupture area

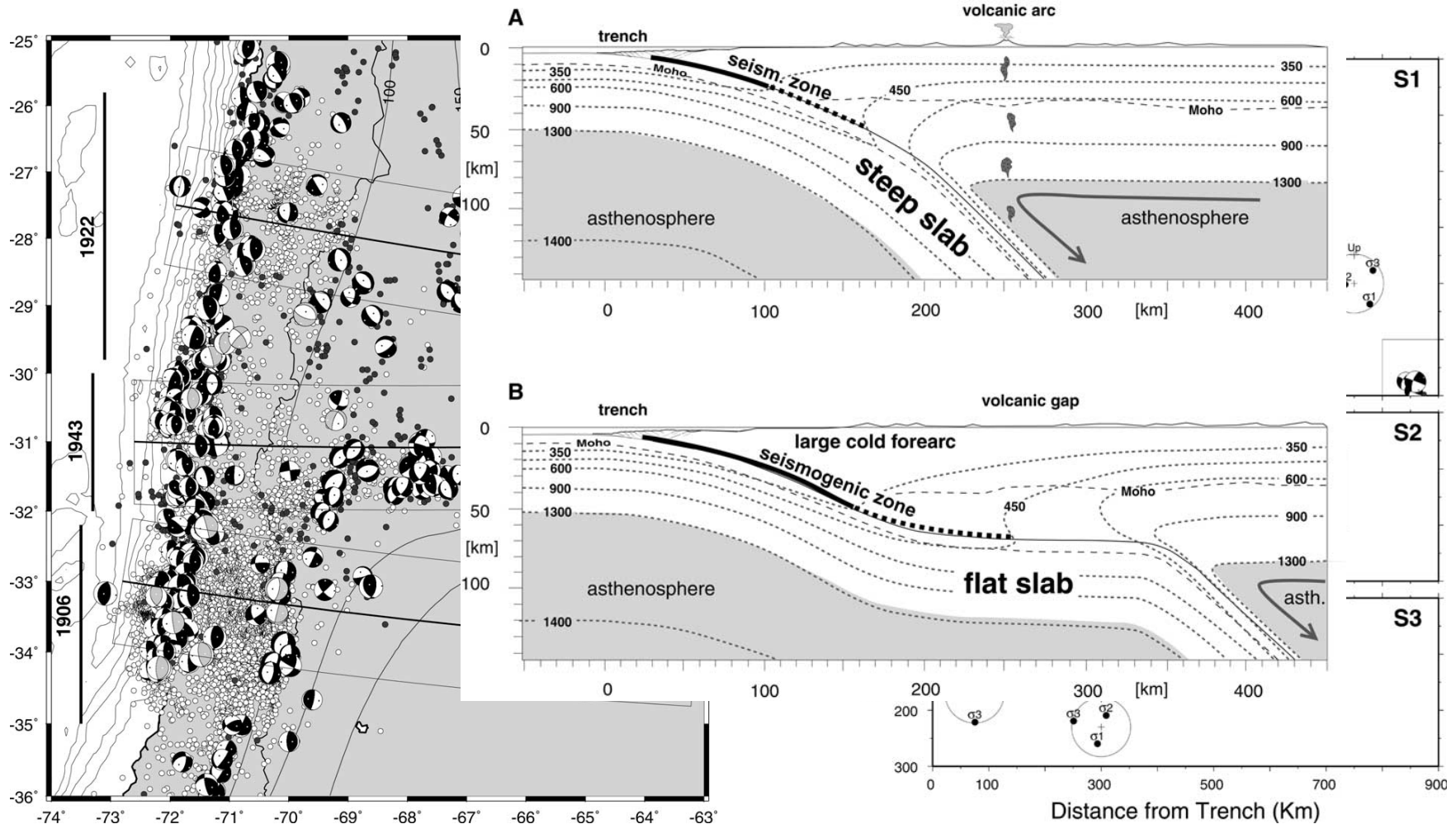


Poli et al., 2017

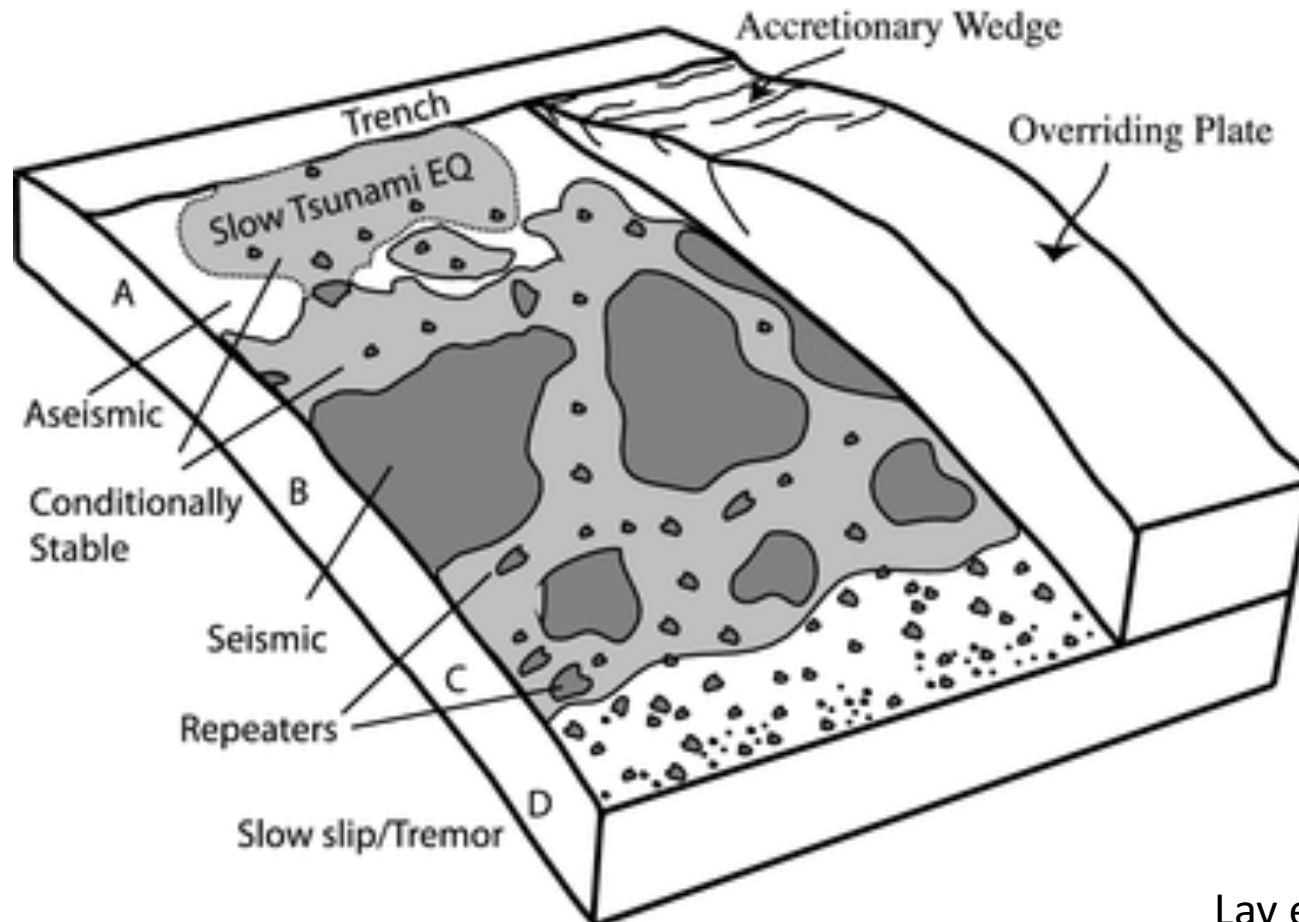
Seismicity cross section: Flat subduction



Seismicity cross section: Flat subduction



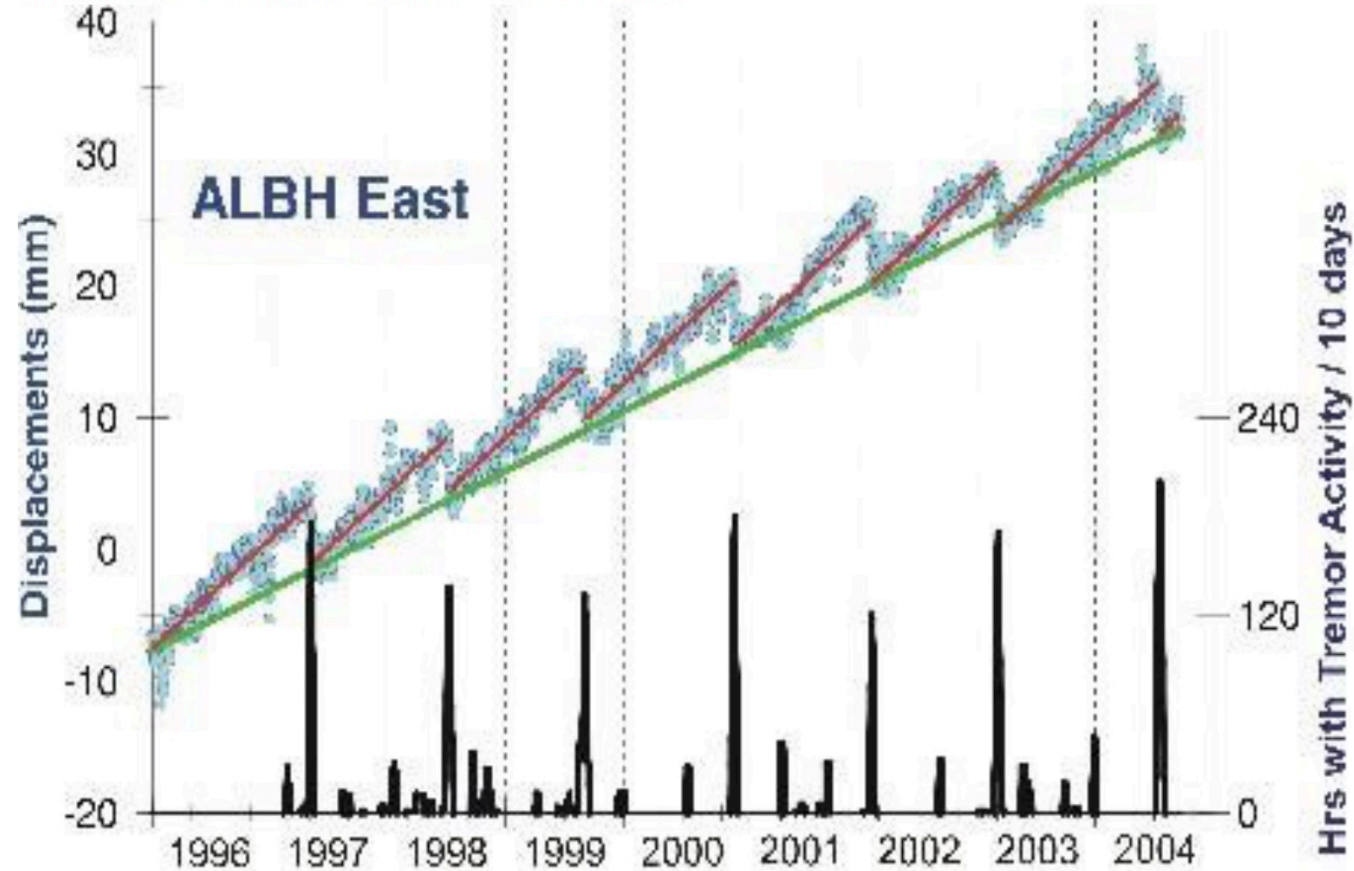
Conceptual model of the subduction zone interface



Lay et al., 2012

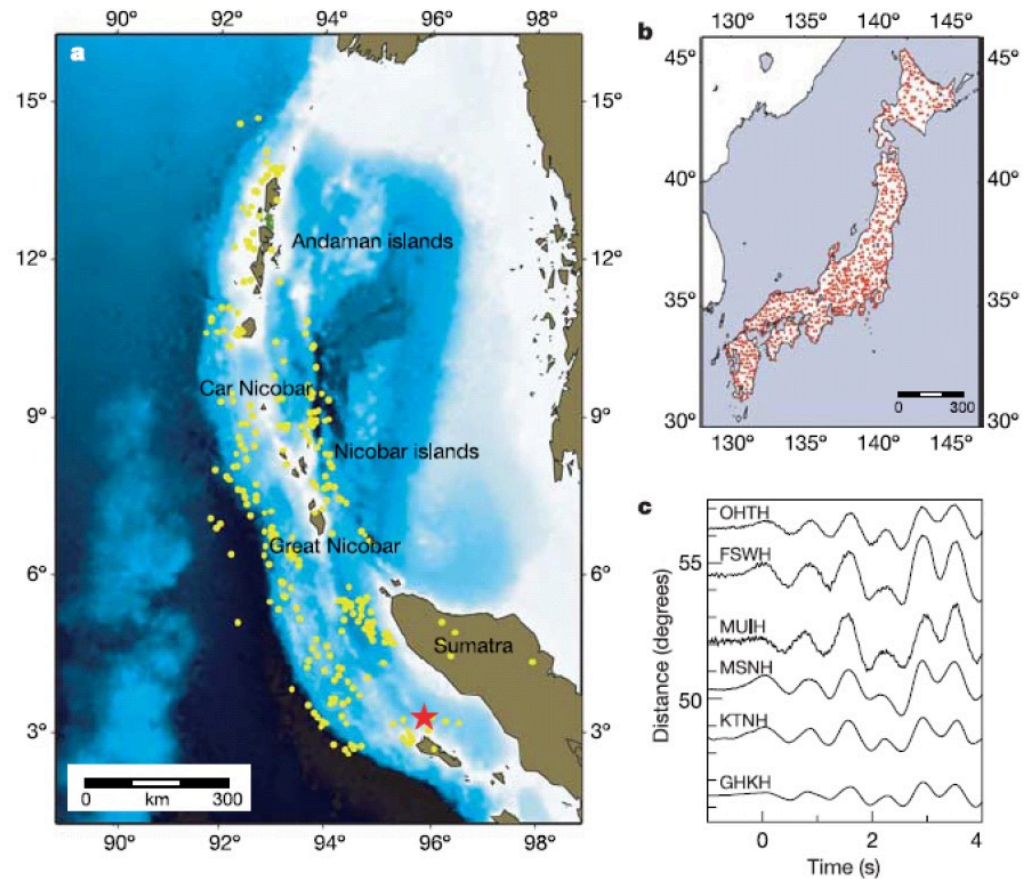
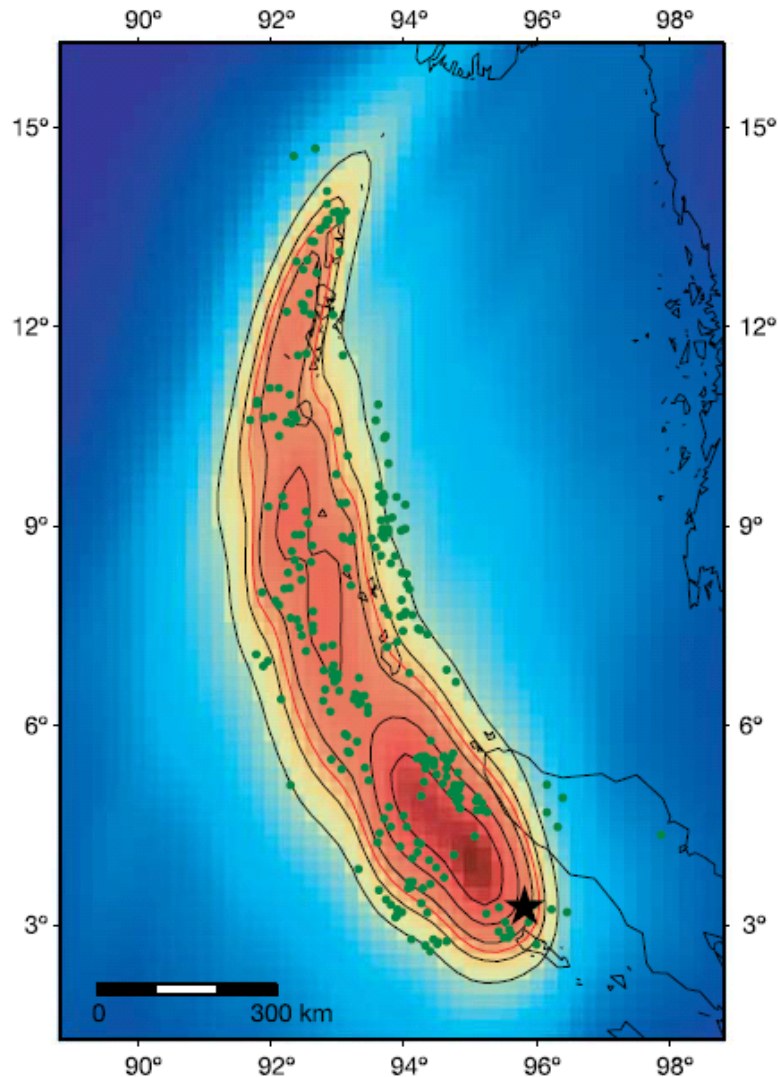
Slow slip events (e.g Cascadia)

Tremor Counts with GPS Data



*source: http://gsc.nrcan.gc.ca/geodyn/ets_e.php

Rupture front imaging

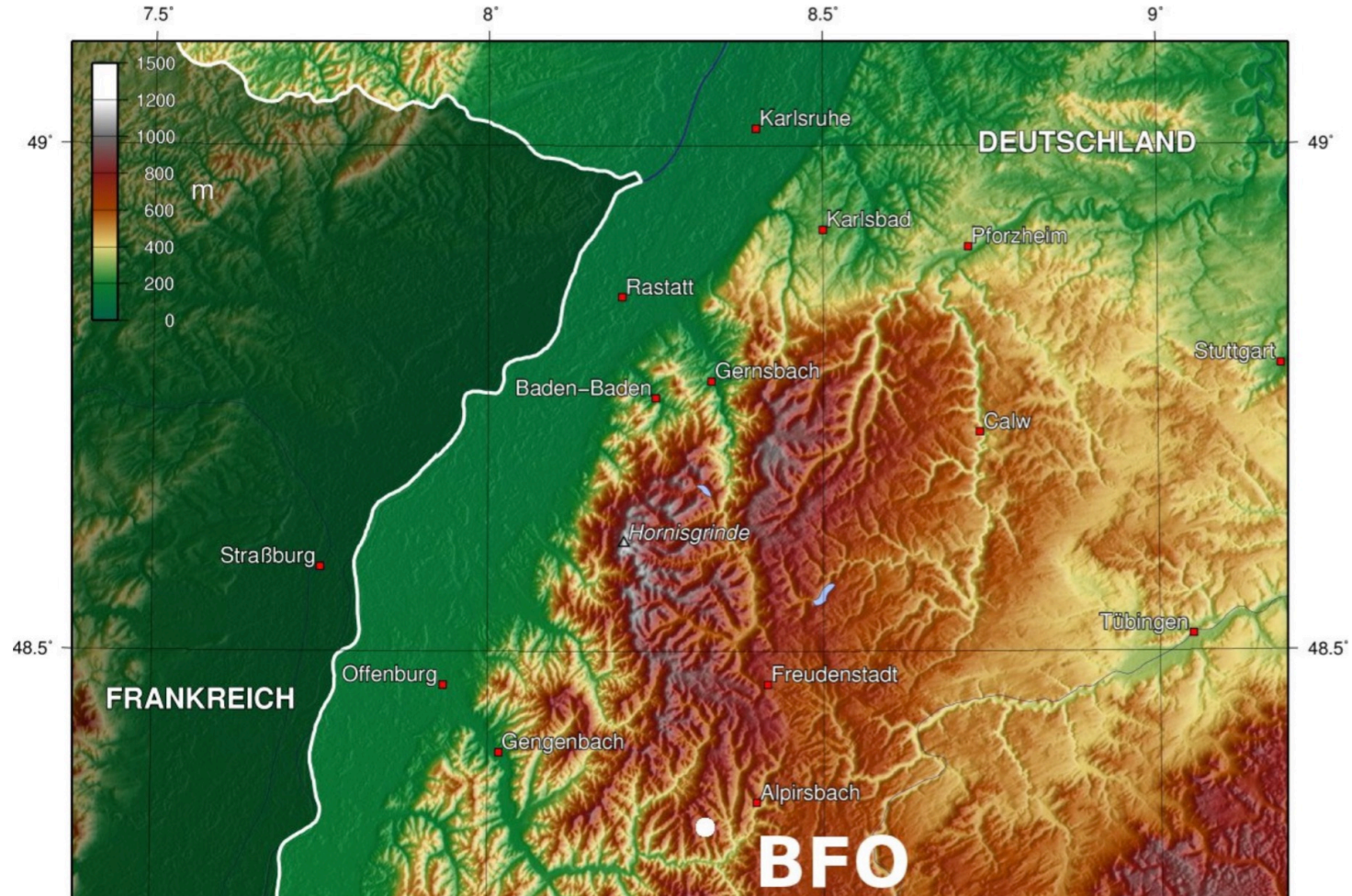


2004 Sumatra earthquake (Ishii et al, 2005)

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BFO observatory



(courtesy of C. Blood, 2006, wikimedia commons)

BFO: Floor Map

Floor map of the gallery

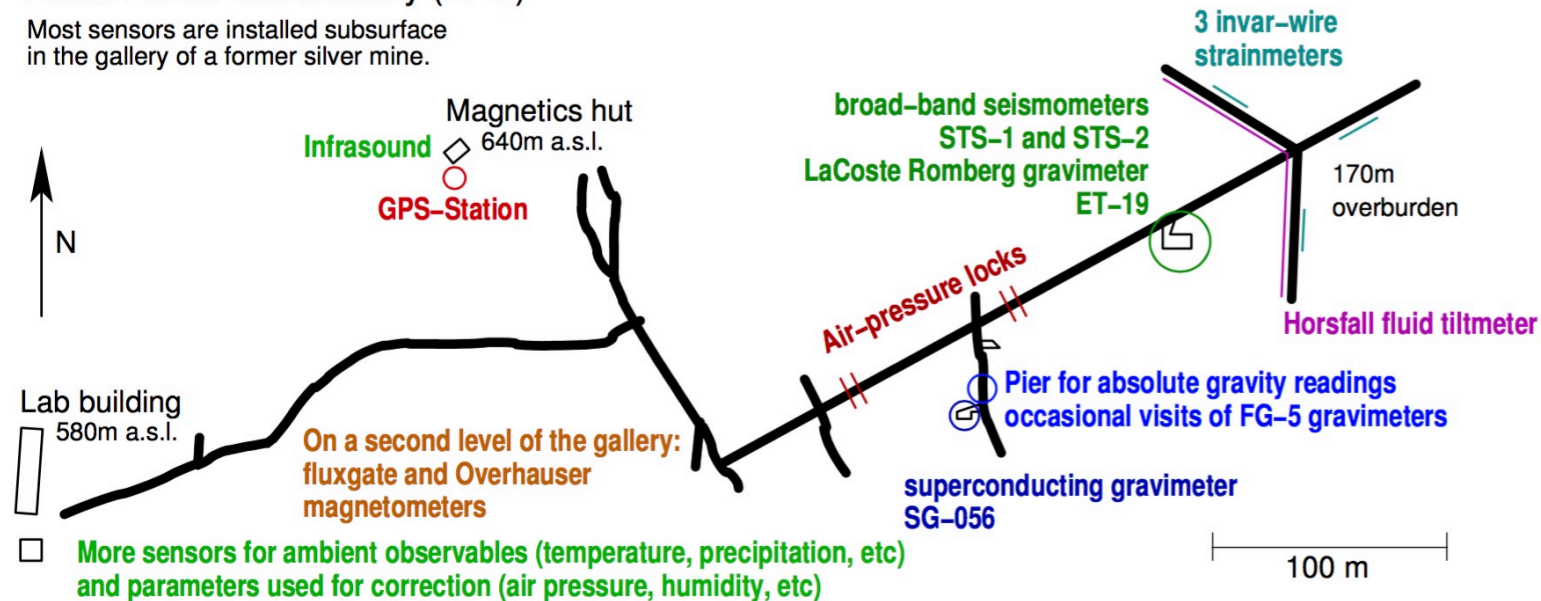


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Black Forest Observatory (BFO)

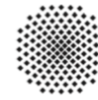
Most sensors are installed subsurface
in the gallery of a former silver mine.



Observables: inertial acceleration, tilt, strain, gravity, magnetic field, position (GPS), ambient pressure, etc.

Approx. 30 sensors, 10 data acquisition systems, observatory clocks, computer networks, emergency power supply system, etc.

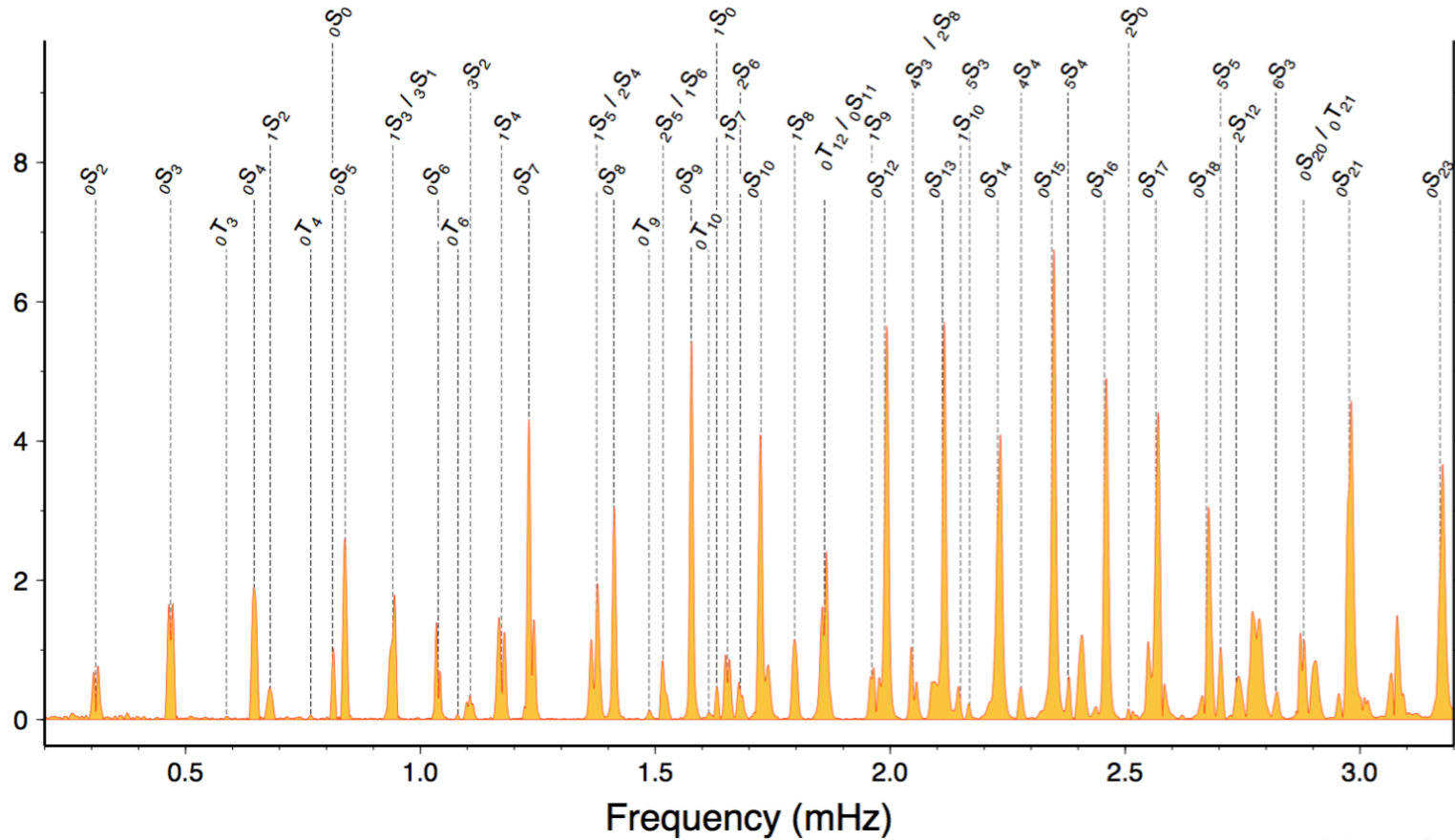
Tohoku quake (Mw 9.0, 11.3.2011)



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Sendai event recorded by SG-056 at BFO (100 hrs)



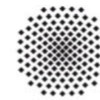
(Widmer-Schnidrig, 2011)

► $1S_1^0C$ ► $0S_2^0$ ► $0S_2^2C, \cos$ ► $0S_3^0$ ► $0S_3^3\cos$ ► $0S_0^0$ ► $0S_0^0C$ ► $0S_8^5\sin$ ► $0T_2^0$ ► $0T_2^0C$ ► $0T_6^6\sin$

(Lucien Saviot, 2011)

Global low noise model

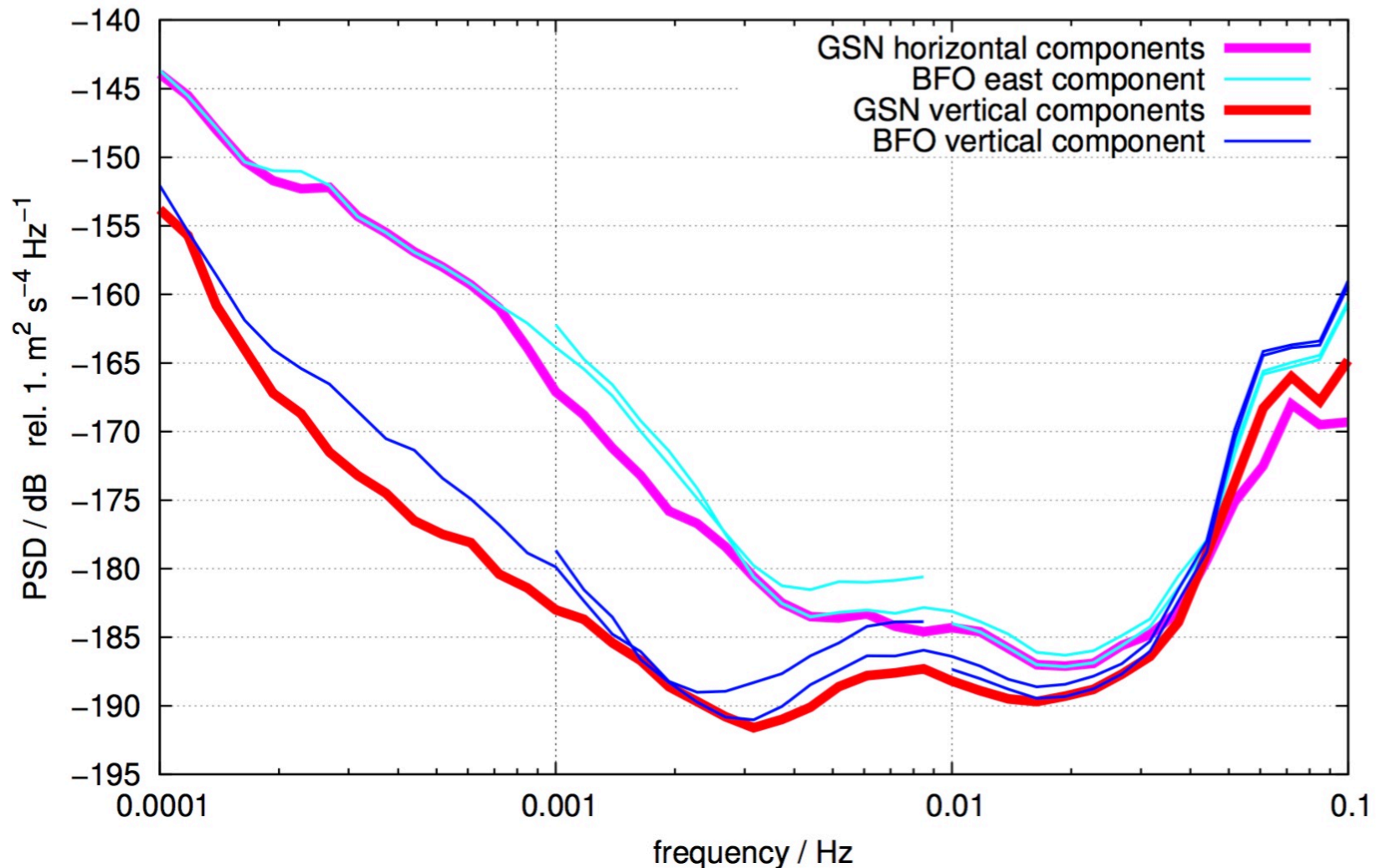
BFO contributes to low noise model



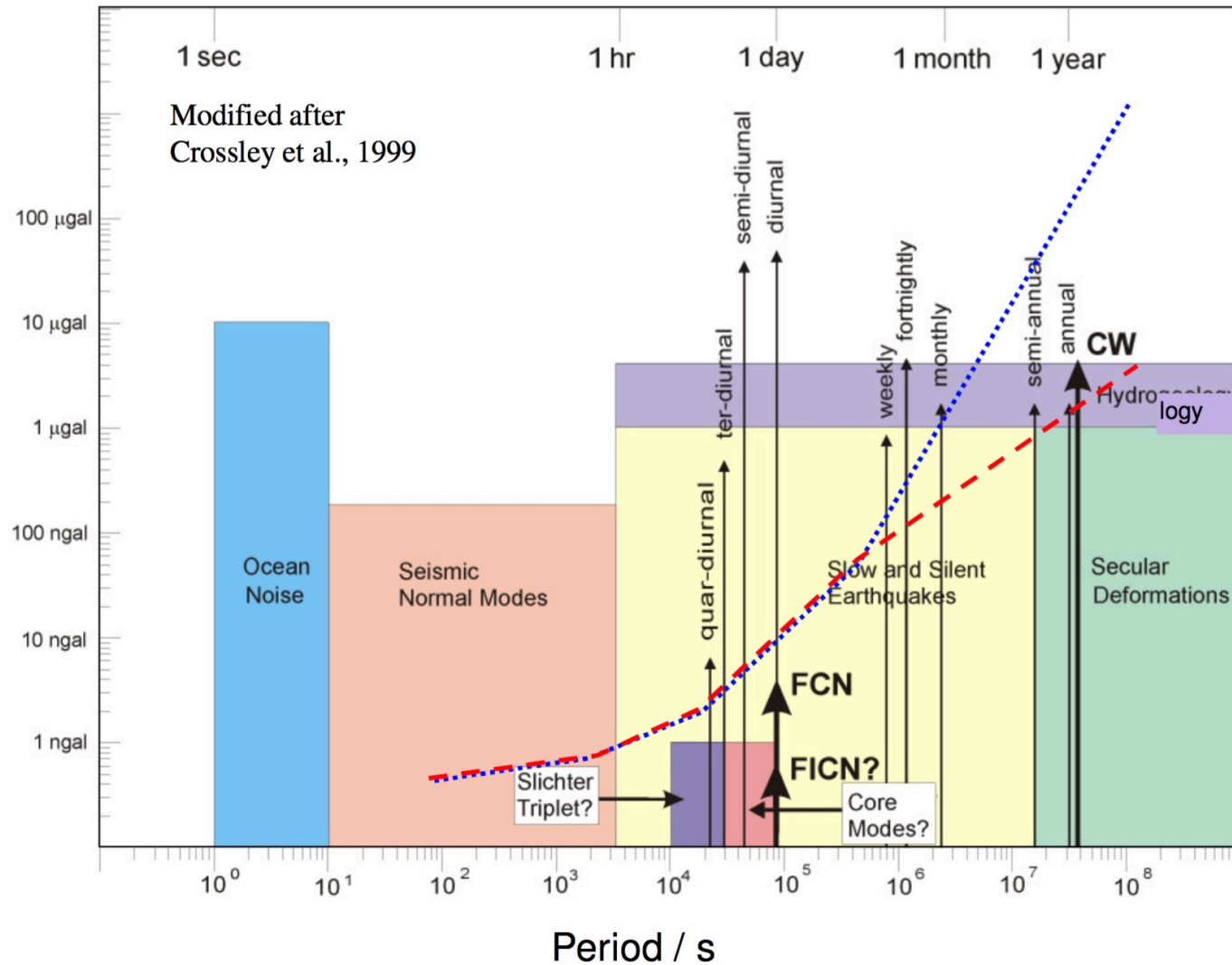
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Low noise model by Berger, Davis, and Ekström (2004).
1st percentile of power spectral density (7/2001 – 6/2002)



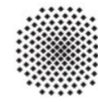
Long period signals



Superconducting gravimeter



Superconducting gravimeter



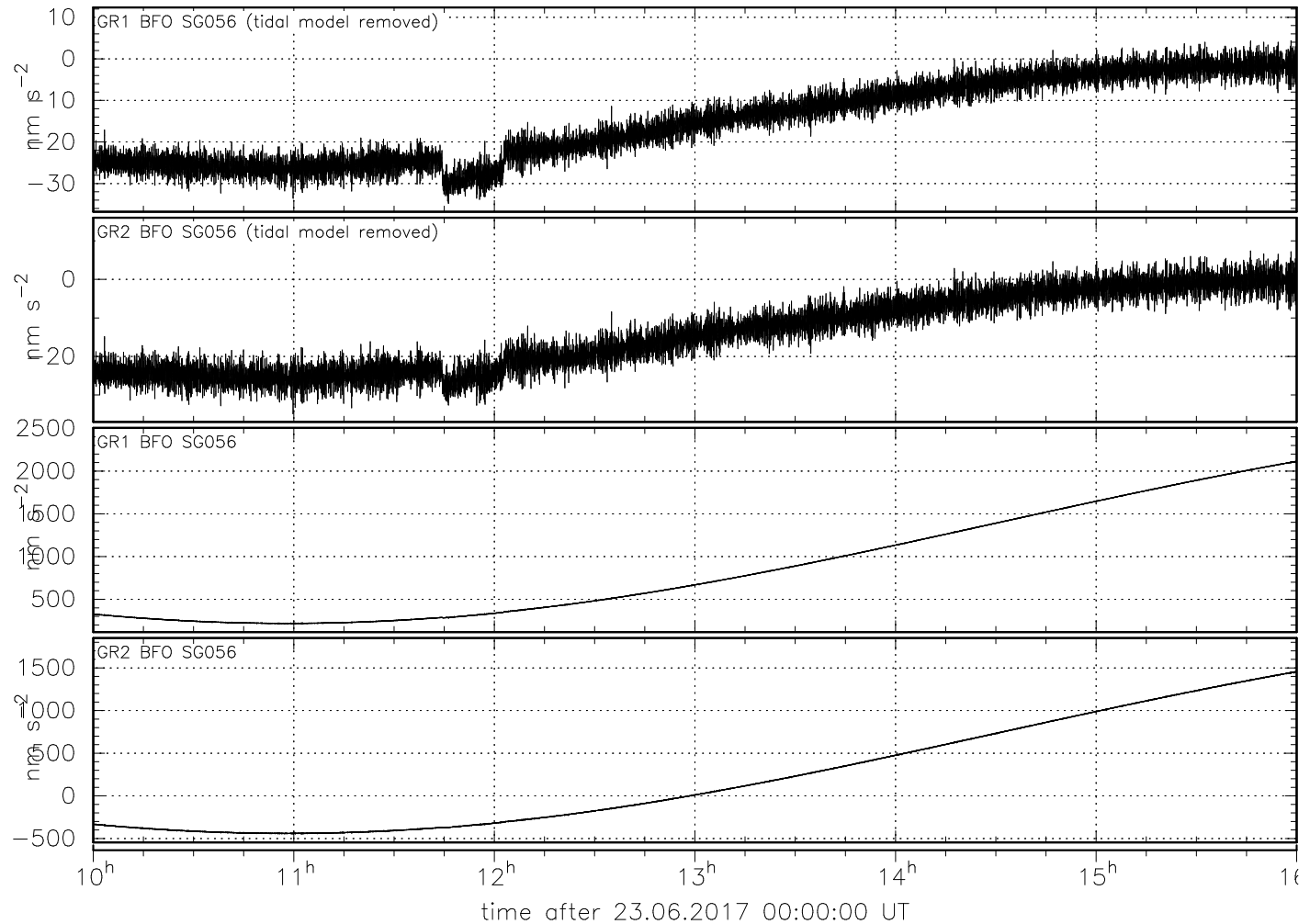
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(courtesy of Markus Breig, 2015)

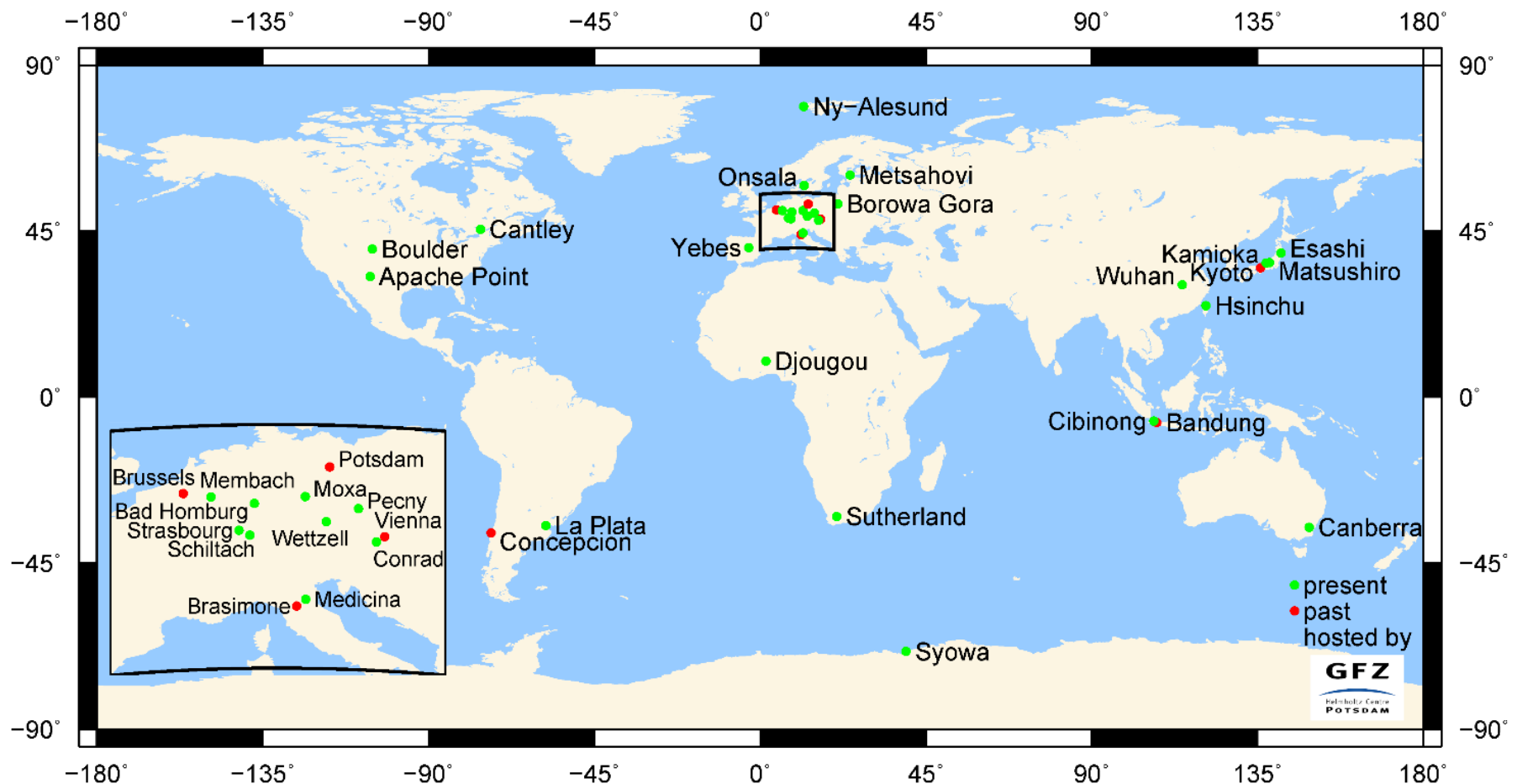
Sensitivity: human anomaly....

Visit of 5 persons in SGK (approx. 11:45 – 12:05 UT)



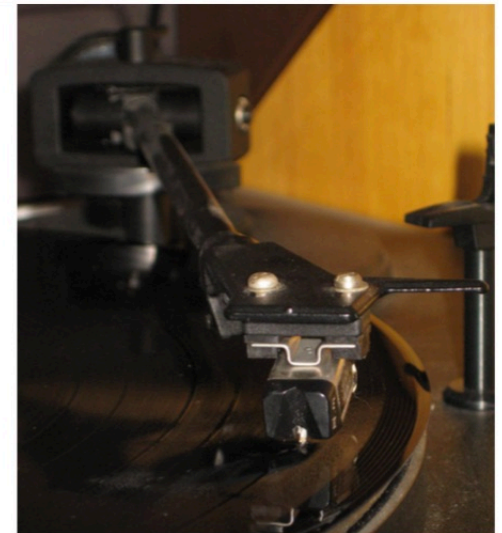
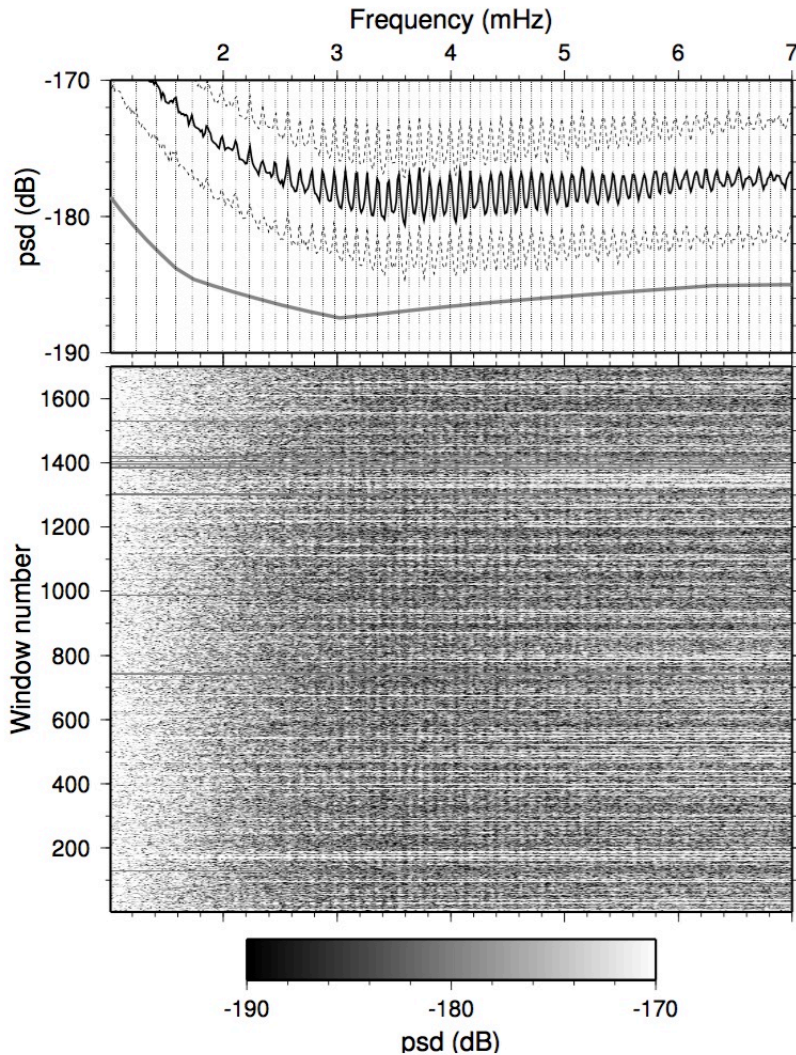
Superconducting gravimeters world wide

IGETS data base containing data from 35 stations



Background free oscillations (Hum of the Earth)

Signalamplitude: $10 \text{ pm s}^{-2} = 10^{-12} \text{ g}$



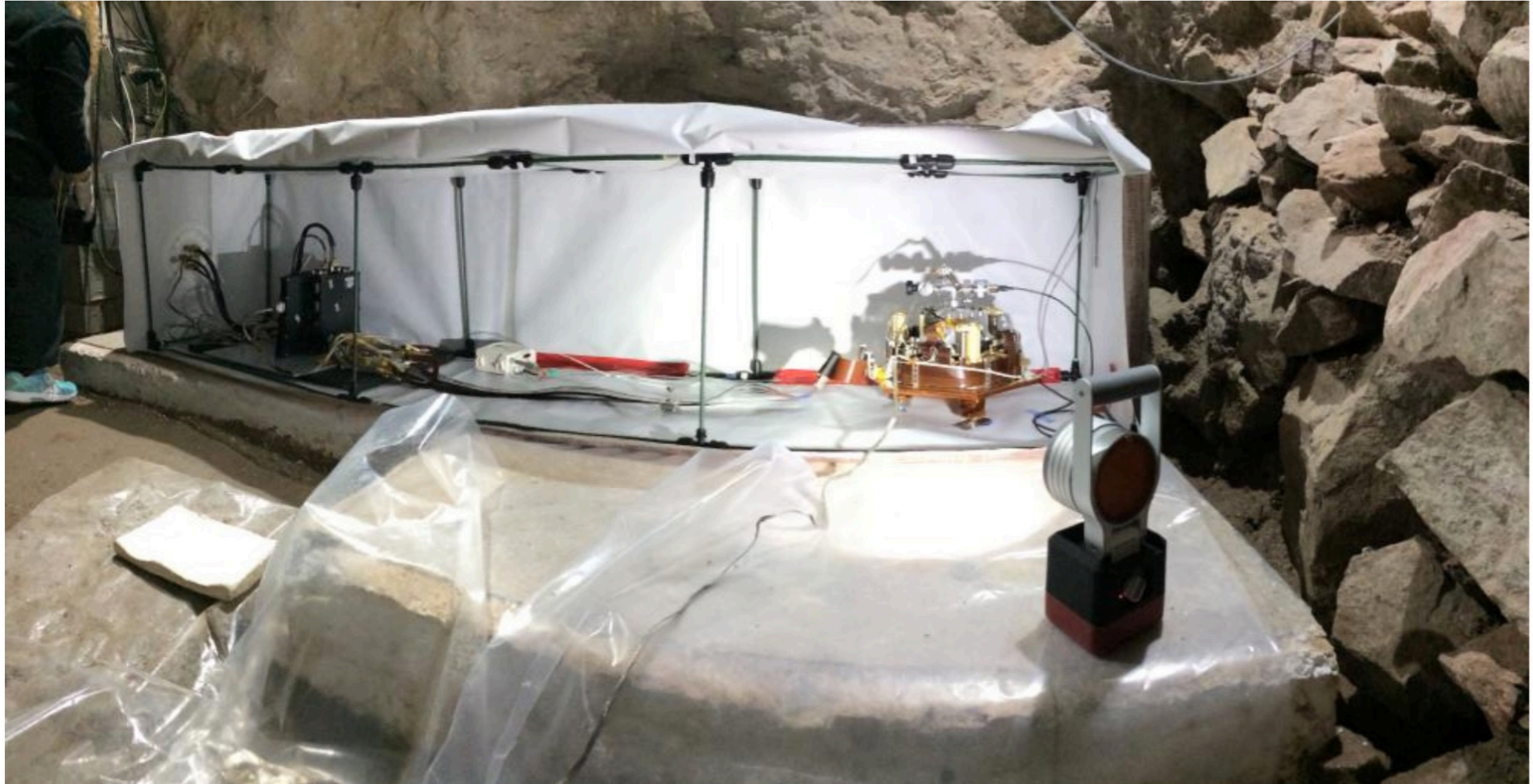
(Wielandt und Widmer-Schnidrig, 2002)

Testing Mars-seismometers

Guests from the french space agency (CNES)



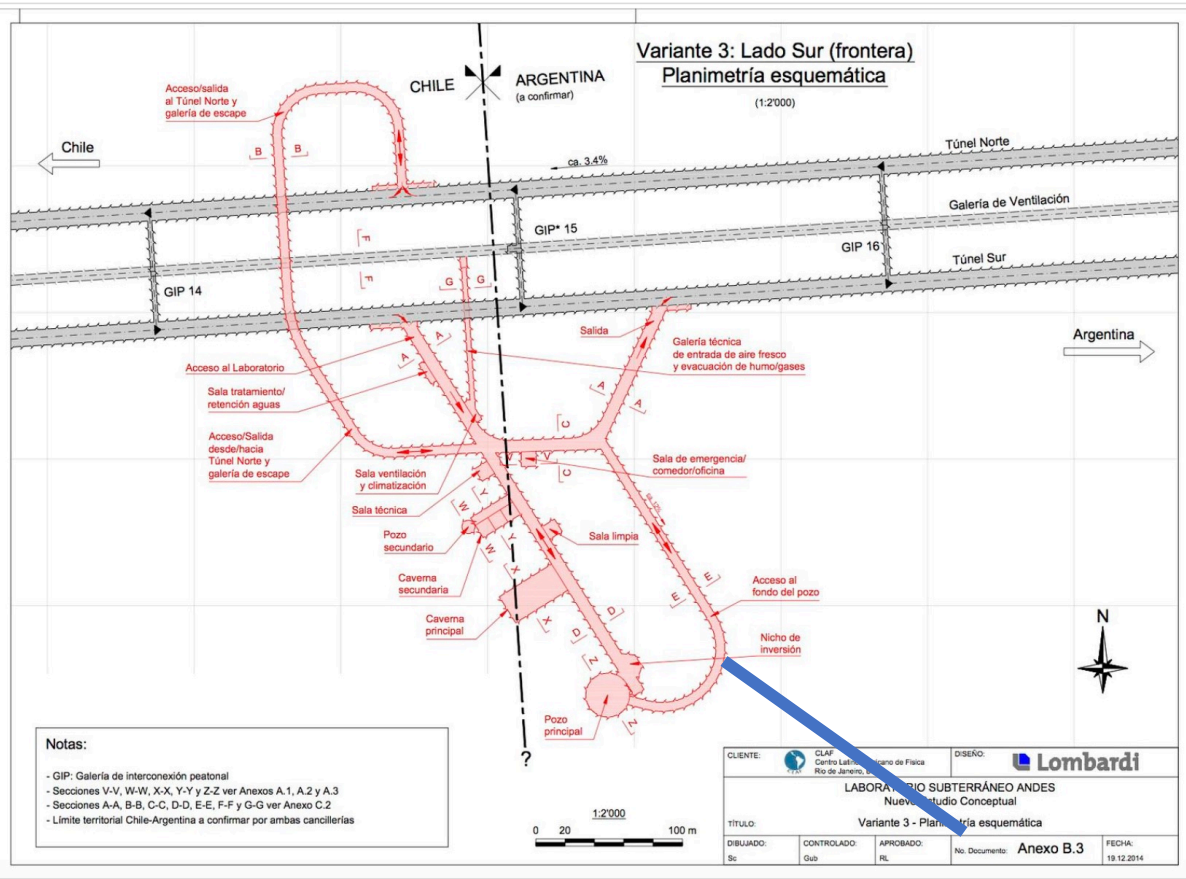
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Laboratory design

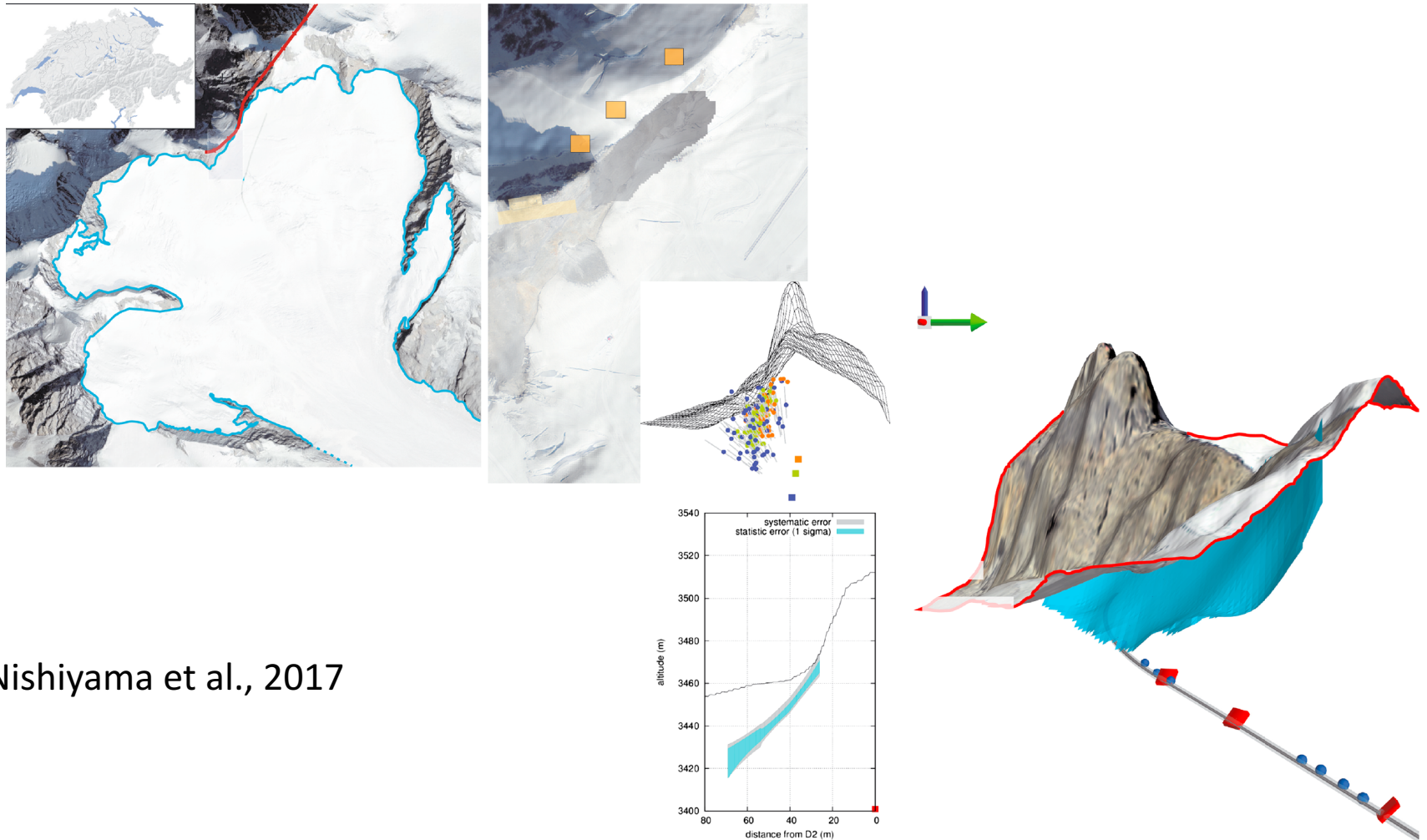


- Small tunnel (access by foot)
- 2 airlocks to separate permanent experiments from visiting experiments
- No or reduced active ventilation
- About 6-8 experiment bays (5x5m and 2-3m high); power supply, optic fibre, GPS time provision

Timeline

- Deploy a BB seismic array/antenna (2017/18)
 - Structural constrain on crust and upper mantle
 - Establish seismic hazard at surface and explore spectral ordinates
- Install instruments during the tunnel excavation to record seismic acceleration in the tunnel
 - ➔ provide design spectra for the instrumentation of the experiments
- Permanent instrumentation of the lab:
 - Long Period seismometer (STS1)
 - Superconducting gravimeter (best also to have a station in the lab in Chile and Argentina along the transect)

First measurement of ice-bedrock interface of alpine glaciers by cosmic muon radiography “Imaging density (contrasts)”



Nishiyama et al., 2017

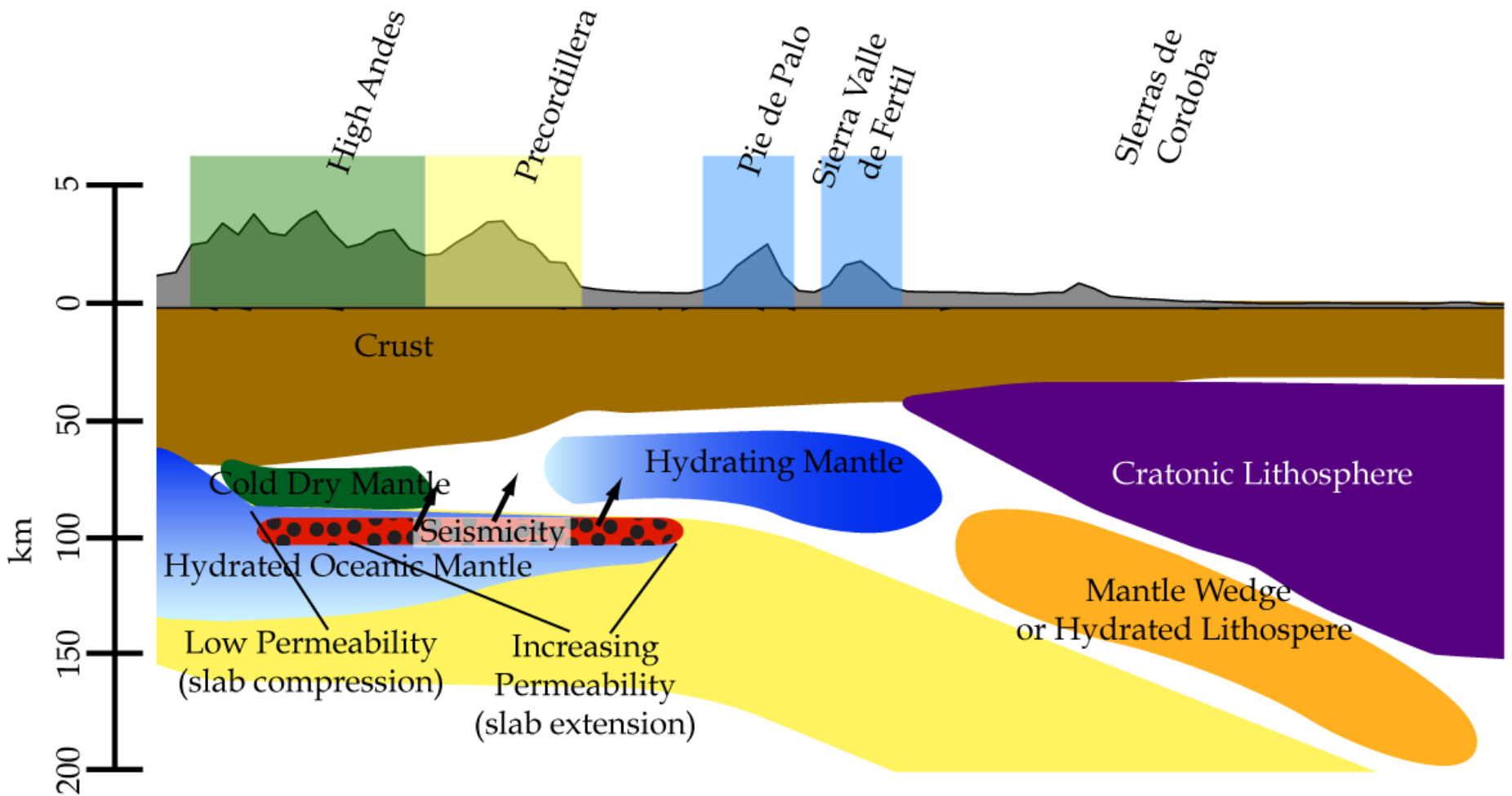


Figure 17